# 22.615, MDH Theory of Fusion Systems Prof. Freidberg

# Lecture 16: Variational Principle

# Variational Principle

$$\omega^2 = \frac{\delta W}{K}$$

$$\delta W = -\frac{1}{2} \int \underline{\xi}^* \cdot \underline{F}(\underline{\xi}) d\underline{r}$$

$$K = \frac{1}{2} \int \rho \left| \xi \right|^2 dr$$

# Advantages:

- 1. allows use of trial functions to estimate  $\omega^2$
- 2. can be applied to multidimensional systems efficiently

### Disadvantages:

- 1. still somewhat complicated
- 2. gives more information than minimum required

#### **Energy Principle**

- 1. Sometimes we only want to know whether the system is stable or not
- 2. No great need to know eigenvalues  $\omega^2$
- 3. Growth rate are very fast  $r^2 \sim v_T^2/a^2 \sim (10 \mu sec)$
- 4. Experimental times ~ 10 msec sec's.
- 5. Since MHD instabilities are very strong, it is more important to know whether system is stable or not, rather than know the precise growth rate (which can be easily estimated)
- 6. In these cases, the variational principle can be simplified further, yielding the Energy Principle. This is a simpler variational procedure which accurately gives <u>stability boundaries</u> but only estimates growth rates.

### The Energy Principle

1. Variational Principle 
$$\omega^2 = \frac{\delta W}{K}$$

If all  $\omega^2 > 0$ , the system is stable

2. Energy Principle  $\delta W \ge 0$  for all allowable displacements, the system is stable. Any displacement which makes  $\delta W < 0 \Rightarrow$  instability

Proof: (based on normal modes) more general proof in text book

1. Assume complete set of normal modes, orthonormal

$$-\omega_n^2 \rho \underline{\xi}_n = \underline{F}\left(\underline{\xi}_n\right) \qquad \int \rho \underline{\xi}_H^* \cdot \xi_{\underline{m}} d\underline{r} = \xi_{mn}$$

2. Arbitrary trial function

$$\underline{\xi} = \sum a_n \, \underline{\xi}_n$$

3. Evaluate δW

$$\begin{split} \delta W &= -\frac{1}{2} \int \underline{\xi}^{\star} \cdot \underline{F} \left( \underline{\xi} \right) = -\frac{1}{2} \sum a_n^{\star} a_m \int \underline{\xi}_n^{\star} \cdot \underline{F} \left( \underline{\xi}_m \right) d\underline{r} \\ &= -\frac{1}{2} \sum a_n^{\star} a_m \int \underline{\xi}_n^{\star} \cdot \left( -\omega_m^2 \rho \underline{\xi}_m \right) \\ &= \frac{1}{2} \sum \omega_m^2 \left| a_m \right|^2 \end{split}$$

- 4. If a trial function can be found which makes  $\delta W<0$  , then at least one  $\omega_m^2<0\to\mbox{ instability}$
- 5. If all trial function make  $\delta W > 0$ , then all  $\omega_m^2 > 0 \rightarrow \text{ stability}$

## **Extended Energy Principle**

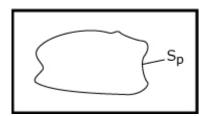
$$\delta W = -\frac{1}{2} \int \underline{\xi}^* \cdot \underline{F} \left( \underline{\xi} \right) d\underline{r}$$

$$\underline{F}\left(\underline{\xi}\right) = \underline{J}_1 \times \underline{B}_0 + \underline{J}_0 \times \underline{B}_1 - \nabla p_1$$

$$= \left[ \nabla \times \nabla \times \left( \underline{\xi} \times \underline{B} \right) \right] \times \underline{B} + \left( \nabla \times \underline{B} \right) \times \left[ \nabla \times \left( \underline{\xi} \times \underline{B} \right) \right] \times \nabla \left[ \underline{\xi} \cdot \nabla p + rp\nabla \cdot \underline{\xi} \right]$$

1. Valid with wall on plasma

$$\underline{\mathbf{n}} \cdot \underline{\boldsymbol{\xi}} \Big|_{S_p} = 0$$

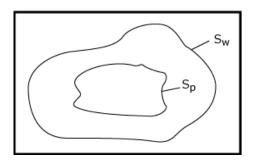


2. Valid with vacuum region

$$[\![\underline{n} \cdot \underline{B}]\!]_{S_n} = 0$$

$$\left\| p + \frac{B^2}{2\mu_0} \right\|_{S_p} = 0$$

$$\underline{\mathbf{n}} \cdot \underline{\mathbf{B}}_1 \big|_{\mathbf{S}_{w}} = \mathbf{0}$$



- 3.  $\delta W$  above not convenient because of complicated boundary condition with wall, and no explicit appearance of Vacuum energy.
- 4. These are resolved by Extended Energy Principle

### **Extended Energy Principle**

1. Rewrite δW<sub>1</sub> introduce natural boundary conditions

$$2. \quad \delta W = -\frac{1}{2}\int d\underline{r} \Big[ \nabla \times \nabla \times \left(\underline{\xi} \times \underline{B}\right) \Big] \times \frac{\underline{B}}{\mu_0} + \frac{\nabla \times B}{\mu_0} \times \Big[ \nabla \times \xi \times B \Big] + \nabla \left(\underline{\xi} \cdot \nabla p + rp\nabla \cdot \underline{\xi}\right) \Big] \cdot \underline{\xi}^*$$
 integrate by parts integrate by parts

3. Define  $\underline{Q} \equiv \underline{B}_1 = \nabla \times (\xi \times \underline{B})$ 

$$\delta W = + \frac{1}{2} \int d\underline{r} \left\{ \frac{\left|\underline{Q}\right|^2}{\mu_0} + rp \left|\nabla \cdot \underline{\xi}\right|^2 - \underline{\xi}^{\star} \cdot \left[\underline{J} \times \underline{Q} + \nabla \left(\underline{\xi} \cdot \nabla p\right)\right] \right\} - \frac{1}{2} \int ds \left(\underline{n} \cdot \underline{\xi}^{\star}\right) \left[rp \nabla \cdot \underline{\xi} - \underline{\underline{B} \cdot \underline{B}_1}{\mu_0}\right]$$

- 4. Separate  $\xi_{\perp}$ ,  $\xi_{\parallel}$ :  $\underline{\xi} = \underline{\xi_{\perp}} + \xi_{\parallel} \underline{b}$
- 5. It is easily shown that  $\underline{b} \cdot \left[ \underline{J} \times \underline{Q} + \nabla \left( \underline{\xi} \cdot \nabla p \right) \right] = 0$  so that last term becomes

$$\underbrace{\xi_{\perp}^{\star}}_{} \cdot \left[ \underline{J} \times \underline{Q} + \nabla \left( \underline{\xi} \cdot \nabla p \right) \right]$$
 integrate by parts

note 
$$\underline{Q} = \nabla \times (\underline{\xi} \times \underline{B}) = \nabla \times (\underline{\xi_{\perp}} \times \underline{B})$$
 
$$\underline{\xi} \cdot \nabla p = \xi_{\perp} \cdot \nabla p$$

6.  $\delta W = \delta W_F + B.T.$ 

$$\delta W_{F} = \frac{1}{2} \int d\underline{r} \left[ \frac{\left|\underline{Q}\right|^{2}}{\mu_{0}} - \underline{\xi_{\perp}^{*}} \cdot \underline{J} \times \underline{Q} + rp \left|\nabla \cdot \underline{\xi}\right|^{2} + \left|\underline{\xi_{\perp}} \cdot \nabla p \right| \nabla \cdot \underline{\xi_{\perp}^{*}} \right]$$

Standard form of the fluid energy

$$BT = \frac{1}{2} \int \! dS \ \underline{n} \cdot \underline{\xi_{\perp}^{*}} \left[ \frac{\underline{B} \cdot \underline{B}_{1}}{\mu_{0}} - rp \, \nabla \cdot \underline{\xi} - \underline{\xi_{\perp}} \cdot \nabla p \, \right]$$

7. Introduce natural boundary condition

$$\left[ p + \frac{B^2}{2\mu_0} \right] = 0 \qquad \text{linearize}$$

$$\begin{bmatrix} p_1 + \frac{B \cdot B_1}{\mu_0} + \underline{\xi} \cdot \nabla \Bigg( p + \frac{B^2}{2\mu_0} \Bigg) \bigg]_{S_p} = \Bigg[ \frac{\widehat{B} \cdot \widehat{B}_1}{\mu_0} + \underline{\xi} \cdot \nabla \frac{\widehat{B}^2}{2\mu_0} \Bigg] \\ -rp\nabla \cdot \underline{\xi} - \xi_\perp \cdot \nabla p$$

8. Substitute above

$$BT = \delta W_s + \frac{1}{2} \int \underline{n} \cdot \underline{\xi}^* \frac{\left(\underline{\hat{B}} \cdot \underline{\hat{B}}_1\right)}{\mu_0} dS$$

$$=\frac{1}{2}\int dS \left|\underline{n}\cdot\underline{\xi}\right|^2 n\cdot \left[\!\!\left[\nabla\!\left(p+\frac{B^2}{2\mu_0}\right)\!\right]\!\!\right]$$

Surface term is non-zero only if surface currents flow

$$9. \quad \delta W = \delta W_F + \delta W_S + \frac{1}{2} \int dS \, \underline{n} \cdot \underline{\xi}^* \, \frac{\underline{\hat{B}} \cdot \underline{\hat{B}}_1}{\mu_0}$$

10. Show last term is related to vacuum energy

$$\begin{split} 11.~\delta W_{_{\boldsymbol{V}}} &= \frac{1}{2\mu_{_{\boldsymbol{0}}}} \int_{_{\boldsymbol{V}}} \left| \underline{\boldsymbol{B}}_{1}^{2} \right| d\underline{\boldsymbol{r}} = \frac{1}{2\mu_{_{\boldsymbol{0}}}} \int_{_{\boldsymbol{0}}} \left| \boldsymbol{\nabla} \times \underline{\boldsymbol{A}}_{1} \right|^{2} d\underline{\boldsymbol{r}} \\ &= \frac{1}{2\mu_{_{\boldsymbol{0}}}} \int_{_{\boldsymbol{0}}} d\underline{\boldsymbol{r}} \left[ \boldsymbol{\nabla} \cdot \left( \underline{\boldsymbol{A}}_{1}^{*} \times \boldsymbol{\nabla} \times \underline{\boldsymbol{A}}_{1} \right) - \underline{\boldsymbol{A}}_{1}^{*} \cdot \boldsymbol{\nabla} \overset{\parallel}{\times} \boldsymbol{\nabla} \times \underline{\boldsymbol{A}}_{1} \right] \end{split}$$

$$= -\frac{1}{2\mu_0} \int\limits_S dS \, \underline{n} \cdot \left( \underline{\widehat{A}}_1^\star \times \underline{\widehat{B}}_1 \right)$$

12. But: since 
$$\underline{n} \cdot \underline{\hat{B}}_1 = \underline{n} \cdot \nabla \times \underline{\hat{A}}_1 = \underline{n} \cdot \nabla \times \underline{\xi} \times \underline{\hat{B}}$$

then 
$$\underline{\widehat{A}}_1 = \xi_{\perp} \times \underline{B} + \nabla \phi$$
 Choose  $\underline{\widehat{B}}_1 \cdot \left(\underline{n} \times \nabla \phi\right) = 0$  as gauge

and 
$$\delta W_v = -\frac{1}{2\mu_0} \int dS \ \underline{n} \cdot \left(\underline{\xi_{\perp}^*} \times \underline{B}\right) \times \underline{\hat{B}}_1$$

$$= \frac{1}{2\mu_0} \int dS \ \underline{n} \cdot \underline{\underline{\beta}}^{\star} \ \underline{\widehat{B}} \cdot \underline{\widehat{B}}_1$$

#### **Extended Energy Principle**

$$\delta W = \delta W_F + \delta W_S + \delta W_V$$

#### **Boundary Conditions on trial functions**

$$\underline{n} \cdot \underline{\hat{B}}_1 \Big|_{S_{\text{out}}} = 0$$

$$\begin{split} \underline{n} \cdot \underline{\hat{B}}_1 \Big|_{S_p} &= \underline{n} \cdot \underline{B}_1 \Big|_{S_p} = \underline{n} \cdot \nabla \times \left(\underline{\xi} \times \underline{B}\right) \Big|_{S_p} \\ &= -\left(\underline{n} \cdot \underline{\xi}\right) \Big[\underline{n} \cdot \left(\underline{n} \cdot \nabla\right) \underline{B}\Big] + \underline{B} \cdot \nabla \left(\underline{n} \cdot \underline{\xi}\right) \Big|_{S_n} \end{split}$$

depends only on  $\underline{n} \cdot \underline{\xi}$ 

pressure balance conditions not required → natural boundary conditions

# **Final Step**

Intuitive form of  $\delta W_{E}$ 

- 1. Standard form OK
- 2. Intuitive form gives more insight.

$$3. \quad \delta W_F \, = \frac{1}{2} \int \underline{\text{d}} r \left\{ \frac{\left|\underline{Q}\right|^2}{\mu_0} - \underline{\xi_\perp^*} - \underline{J} \times \underline{Q} + rp \left|\nabla \cdot \underline{\xi}\right|^2 + \left(\underline{\xi_\perp} \cdot \nabla p\right) \nabla \cdot \underline{\xi_\perp^*} \right\}$$

$$\left|\underline{Q}\right|^2 = \left|Q_{\perp}\right|^2 + \left|Q_{\parallel}\right|^2$$

$$\begin{split} \underline{\xi_{\perp}^{*}} \cdot \underline{J} \times \underline{Q} &= \left(\underline{\xi_{\perp}^{*}} \times \underline{b}\right) \cdot \underline{Q_{\perp}} J_{\parallel} + Q_{\parallel} \underline{\xi_{\perp}^{*}} \cdot \underline{J_{\perp}} \times \underline{b} \\ \\ \text{now:} \qquad \underline{J_{\perp}} &= \underline{\underline{b}} \times \nabla \underline{p} \\ & \underline{Q}_{\parallel} = \underline{\underline{b}} \cdot \nabla \times \left(\underline{\xi_{\perp}} \times \underline{B}\right) \\ \\ &= \underline{\underline{b}} \cdot \left(\underline{B} \cdot \nabla \underline{\xi_{\perp}} - \underline{\xi_{\perp}} \cdot \nabla \underline{B} - \underline{B} \nabla \cdot \underline{\xi_{\perp}}\right) \\ \\ &= -B \left(\nabla \cdot \underline{\xi_{\perp}} + 2\underline{\xi_{\perp}} \cdot \underline{\kappa}\right) + \frac{\mu_{0}}{B} \underline{\xi_{\perp}} \cdot \nabla \underline{p} \end{split}$$

Substitute back

1 2 3

$$\delta W_{F} = \frac{1}{2} \int d\underline{r} \left[ \frac{\left|\underline{Q_{\perp}}\right|^{2}}{\mu_{0}} + \frac{B^{2}}{\mu_{0}} \left| \nabla \cdot \underline{\xi_{\perp}} + 2\underline{\xi_{\perp}} \cdot \underline{\kappa} \right|^{2} \\ + rp \left| \nabla \cdot \underline{\xi} \right|^{2} \\ - 2 \left( \underline{\xi_{\perp}} \cdot \nabla p \right) \left( \underline{\kappa} \cdot \underline{\xi_{\perp}^{*}} \right) - J_{\parallel} \left( \underline{\xi_{\perp}^{*}} \times \underline{b} \right) \cdot \underline{Q_{\perp}} \right]$$

- 1. line bending > 0
- shear alform wave
- 2. magnetic compression > 0 compressional alform wave
- 3. plasma compression > 0
- sound wave
- 4. pressure driven modes + or -
- 5. current driven modes (kinks) + or -

#### **Summary**

Energy Principle:  $\delta W = \delta W_F + \delta W_S + \delta W_V$ 

 $\delta W \ge 0$  for all allowable displacements  $\rightarrow$  stability

 $\delta W < 0$  for any allowable displacement  $\rightarrow$  instability

Minimize  $\delta W$  with respect to three components of  $\xi$  .

#### Incompressibility

1. Because of the simple way in which  $\xi_{\parallel}$  appears in  $\delta W$  , it is possible to minimize once for all with respect to  $\,\xi_{\parallel}\,$  and eliminate it from the calculation.

2. Only appearance of  $\xi_{\parallel}$ 

$$\delta W_{\parallel} = \int d\underline{r} \ rp \left| \nabla \cdot \underline{\xi} \right|^2$$

- 3. Let  $\xi_{\parallel} \rightarrow \xi_{\parallel} + \delta \xi_{\parallel}$
- 4. Vary  $\delta W_{\parallel}$   $\underline{\xi} = \underline{\xi_{\perp}} + \xi_{\parallel} \underline{\underline{B}}_{R}$

$$\delta \left( \delta W_{\parallel} \right) = \int d\underline{r} \ rp \left( \nabla \cdot \underline{\xi} \right) \nabla \cdot \left( \delta \xi_{\parallel} \, \frac{\underline{B}}{B} \right)$$
 integrate by parts

$$= -\!\int d\!\underline{r} \, \frac{\delta \xi_{||}}{B} \, \, \underline{B} \cdot \nabla \left( rp \, \nabla \cdot \underline{\xi} \right)$$

$$= -\!\int d\underline{r} \, \frac{\delta \xi_{||}}{B} \, \, r p \underline{B} \cdot \nabla \left( \nabla \cdot \underline{\xi} \right)$$

5. Several minimizing condition

$$\underline{\mathbf{B}} \cdot \nabla \left( \nabla \cdot \underline{\boldsymbol{\xi}} \right) = \mathbf{0}$$

6. If  $B \cdot \nabla$  is non-singular then

$$\nabla \cdot \underline{\xi} = 0$$
 (obvious)

$$\delta W_{\parallel} = \frac{1}{2} \int r p \left| \nabla \cdot \underline{\xi} \right|^2 \to 0$$

- 7. Two cases where  $\nabla \cdot \underline{\xi}$  cannot be set to zero
- 8. Special symmetry

$$\text{Example: Z pinch } \underline{B} = B_{\theta}\left(r\right)\underline{e}_{\theta} \qquad \underline{\xi} \sim e^{im\theta + ikz}\underline{\xi}\left(r\right)$$

$$\underline{B} \cdot \nabla \frac{\xi_{\parallel}}{B} = \frac{B_{\theta}}{r} \frac{\partial}{\partial \theta} \frac{\xi_{\parallel}}{B_{\theta}} = \frac{im \xi_{\parallel}}{r}$$

For 
$$m = 0$$
  $\underline{B} \cdot \nabla \frac{\xi_{\parallel}}{B} = 0$ 

$$Note: \ \, \nabla \cdot \underline{\xi} = \nabla \cdot \underline{\xi_{\perp}} + \nabla \cdot \frac{\xi_{\parallel}}{B} \underline{B} = \nabla \cdot \underline{\xi_{\perp}} + \underline{B} \cdot \nabla \, \frac{\xi_{\parallel}}{B}$$

In special symmetry  $\nabla \cdot \underline{\xi} = \nabla \cdot \underline{\xi_{\perp}}$  and  $\xi_n$  does not appear. The term  $\operatorname{rp} \left| \nabla \cdot \underline{\xi_{\perp}} \right|^2$  must be maintained for the rest of the minimization.

9. Closed line (periodicity constraints). Choose  $\,\xi_{\parallel}\,$  so  $\,\nabla\cdot\xi=0\,$ 

$$\underline{B} \cdot \nabla \frac{\xi_{\parallel}}{B} = -\nabla \cdot \underline{\xi_{\perp}} = B \frac{\partial \xi_{\parallel} / B}{\partial I}$$

$$\frac{\xi_{\parallel}}{B} = -\int \frac{\nabla \cdot \xi_{\perp}}{B} dI$$

In general  $\frac{\xi_{\parallel}}{B} (I + L) \neq \frac{\xi_{\parallel}}{B} (I) \rightarrow \text{ no periodicity}$ 

Solution

$$\underline{B}\cdot\nabla\quad\nabla\cdot\xi=0$$

$$\therefore \nabla \cdot \underline{\xi} = F(p)$$

homogenous solution

$$\underline{B} \cdot \nabla \frac{\xi_{||}}{R} = -\nabla \cdot \underline{\xi_{\perp}} + F\left(p\right)$$

$$\frac{\xi_{\parallel}}{B} = -\int \frac{\nabla \cdot \xi_{\perp}}{B} \, dI + \int \frac{F\left(p\right) dI}{B} = -\int_{0}^{I} \frac{\nabla \cdot \xi_{\perp}}{B} \, dI + F\left(p\right) \int_{0}^{I} \frac{dI}{B} \, dI + \int_{0}^{I} \frac{dI}{B}$$

In periodicity choose

$$F\left(p\right) = <\nabla\cdot\underline{\xi_{\perp}}> = \frac{\oint\frac{dI}{B}\nabla\cdot\underline{\xi_{\perp}}}{\oint\frac{dI}{B}}$$

Then 
$$\delta W_{\parallel} = \frac{1}{2} \int rp \left| \nabla \cdot \underline{\xi} \right|^2 d\underline{r} = \frac{1}{2} \int rp \, F^2 \, d\underline{r}$$

$$\delta W_{||} = \frac{1}{2} \int d\underline{r} \, r p \, \Big| < \nabla \cdot \underline{\xi_{\perp}} \, > \Big|^2$$

Only a function of  $\xi_{\perp}$ 

# Summary of internal modes in a straight tokamak

- 1.  $m \ge 2$  stable
- 2. m = 1, n = 1 must use for n = 1, requires q(0) > 1 for stability
- 3. internal modes do not limit  $\beta$ , or I (q(a)), but clamp  $q(0) \approx 1$ , by sawtooth oscillations
- 4. To show instability we needed to calculate  $\delta W=e^2\delta W_2+\varepsilon^4\delta W_4$  0 must

#### Consider now external modes

- 1. Vacuum no force free fields
- 2. m=1 Kruskal Shafranov limit
- 3. High m external kinks